

X-ray Energy Spectroscopy (XES).

X-ray fluorescence as an analytical tool for element analysis is based on 3 fundamental parameters:

- A. Specificity:** In determining an x-ray emission energy $E \rightarrow$ certainty of presence of element A
- B. Sensitivity:** Given the presence of element A \rightarrow what is the concentration of A.
- C. Detection limit:** The smallest amount or concentration that may be ascertained.

Characteristic emission, X-ray fluorescence

a) Simplified image of the atom as a spherical shell structure (the N. Bohr model, Fig.1)

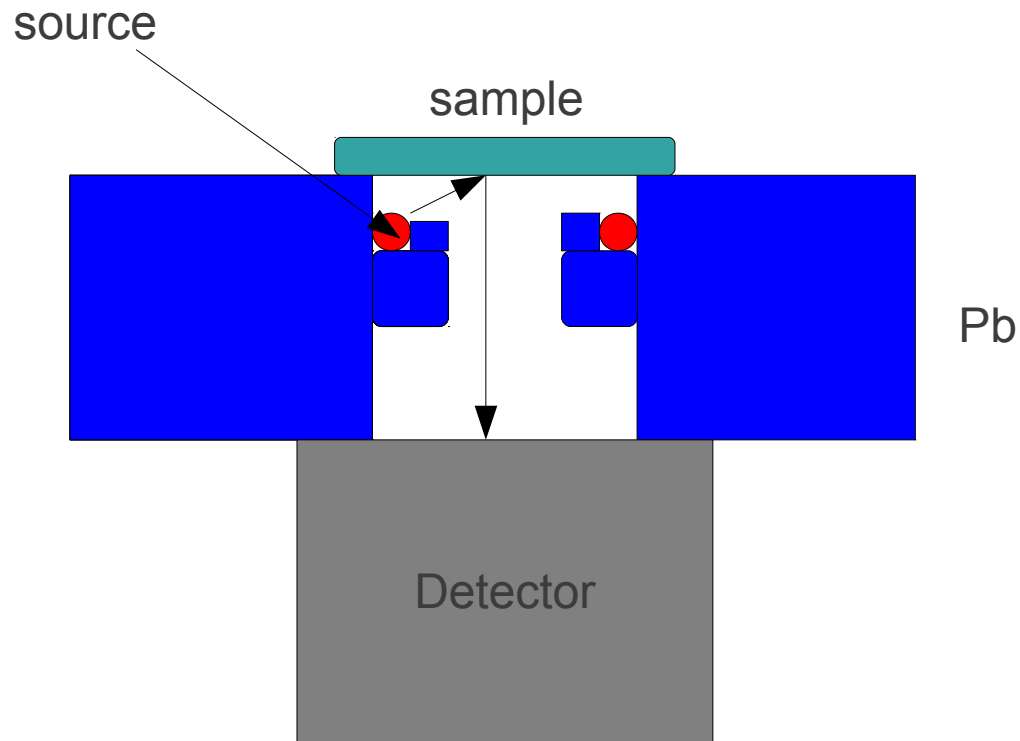
- electrons occupy shells labeled K(n=1), L(n=2)
- M(n=3), N(n=5), O(n=6) etc.; n-principal quantum number

b) Each electron shell has specific binding energy of the electron: K_{ab} , L_{lab} , etc. - absorption edge, the minimum energy required to lift an electron free. Characteristic energy (b.e.) $\leftrightarrow Z$

Cont. of XES

c) Fluorescences, or the generation of secondary radiation from an atom after removal of an electron out of its shell – excitation of the atom → emission of electromagnetic radiation $E_\gamma = E_i - E_f$ where i and f are the initial and final energies of the electron undergoing the transition.

This aspect forms the basis for x-ray energy spectroscopy .



Schematic representation of radioisotope excitation system with direct irradiator

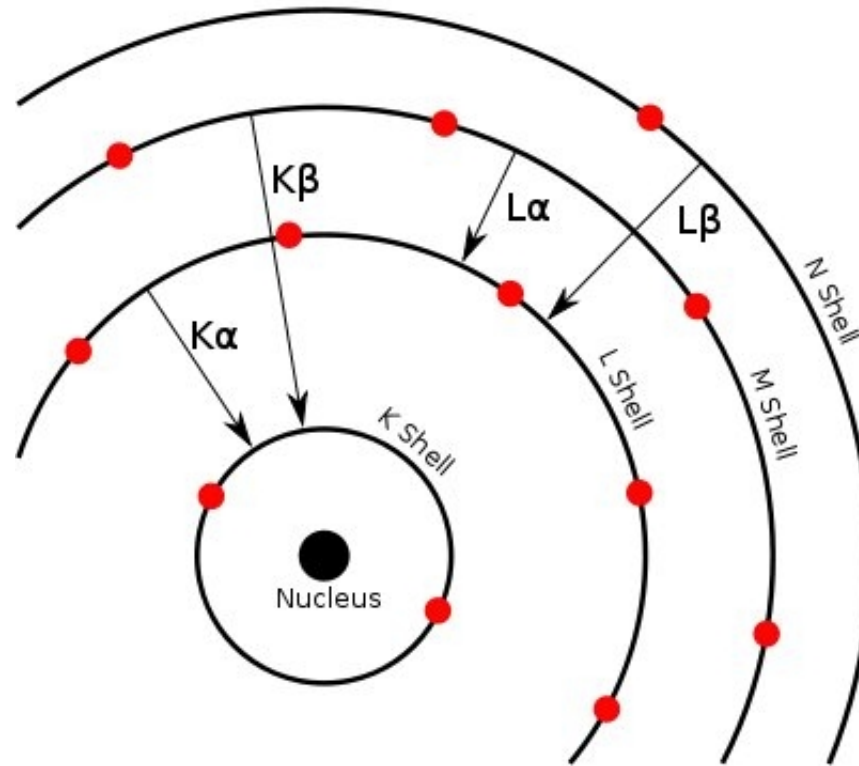
Four basic quantum numbers:

The first, n , describes the **electron shell**, or energy level. The value of ranges from 1 to " n ", where " n " is the shell containing the outermost electron of that atom. For example, in **cesium (Cs)**, the outermost **valence** electron is in the shell with energy level 6, so an electron in cesium can have an value from 1 to 6.

The second, l , describes the **subshell** (0 = s orbital, 1 = p orbital, 2 = d orbital, 3 = f orbital, etc.). The value of l ranges from 0 to $n-1$. This is because the first p orbital ($l=1$) appears in the second electron shell ($n=2$), the first d orbital ($l=2$) appears in the third shell ($n=3$), and so on. A quantum number beginning in 3,0,... describes an electron in the s orbital of the third electron shell of an atom.

The third, m_l , describes the specific **orbital** within that subshell. The values m_l range from $-l$ to l . The s subshell ($l=0$) contain only one orbital, and therefore the m_l of an electron in an s subshell will always be 0. The p subshell ($l=1$) contains three orbitals, so the m_l of an electron in a p subshell will be -1 , 0, or 1. The d subshell: 5 orbitals etc..

The fourth, m_s , describes the spin of the electron within that orbital. Because an orbital never contains more than two electrons, m_s will be either $\frac{1}{2}$ or $-\frac{1}{2}$, corresponding with "spin" and "opposite spin".



Atomic model, showing electron transitions that may follow electron vacancies.

Characteristic X-ray transitions

- K-series

- $K_{\alpha 1} = K - L_{III}$

- $K_{\alpha 2} = K - L_{II}$

- $K_{\beta 1} = K - M_{III}$

- $K_{\beta 3} = K - M_{II}$

- $K_{\beta 2} = K - M_{II,III}$

- L-series

- $L_{\alpha 1} = L_{III} - M_{V}$

- $L_{\alpha 2} = L_{III} - M_{IV}$

- $L_{\beta 1} = L_{II} - M_{IV}$

Radiationless transitions, the Auger effect, the Fluorescence Yield, some complications ..

- This is an internal atomic process that reduces the characteristic x-ray output or yield relative to predictions from **photoelectric** (p-e) cross sections.
- If a K-shell vacancy has been created the x-ray photon that follows is energetic enough to create vacancies by ejecting electrons from higher shells in the same atom (next Fig). In this case, 2 vacancies are created in the L-shell, as secondary internal p-e absorption ejects an L electron – known as an Auger electron. Barring subsequent **Auger processes** in still higher shells, the net result will only be the emission of L and M x-rays.

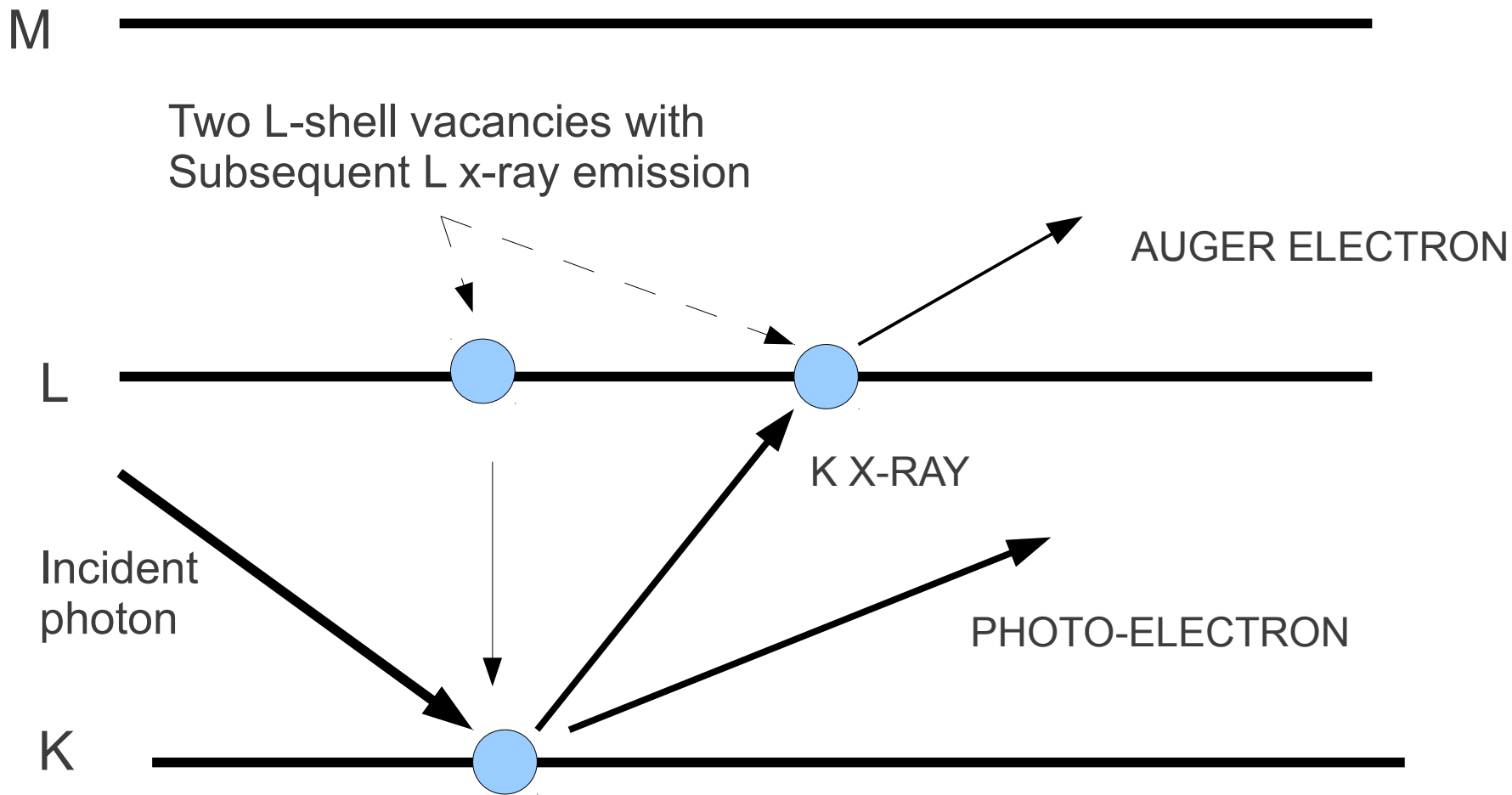
- This competitive effect to characteristic x-ray emission introduces a factor termed **fluorescence yield ω** , which may be defined as the ratio of emitted x-rays to the number of primary vacancies created. It is a function both of atomic number, and of the location (shell) of initial vacancy.

- Another **complication** of XES spectra is **Elastic Scattering** – process where photons of the fluorescing radiation are scattered by atomic electrons that are so tightly bound to the atom that no ionization or excitation is possible. The collision is, therefore, effectively with the atom as an entity – with no energy loss ! The Z dependence of elastic scatter is approximately as Z^2 in the region of interest for XES analysis.

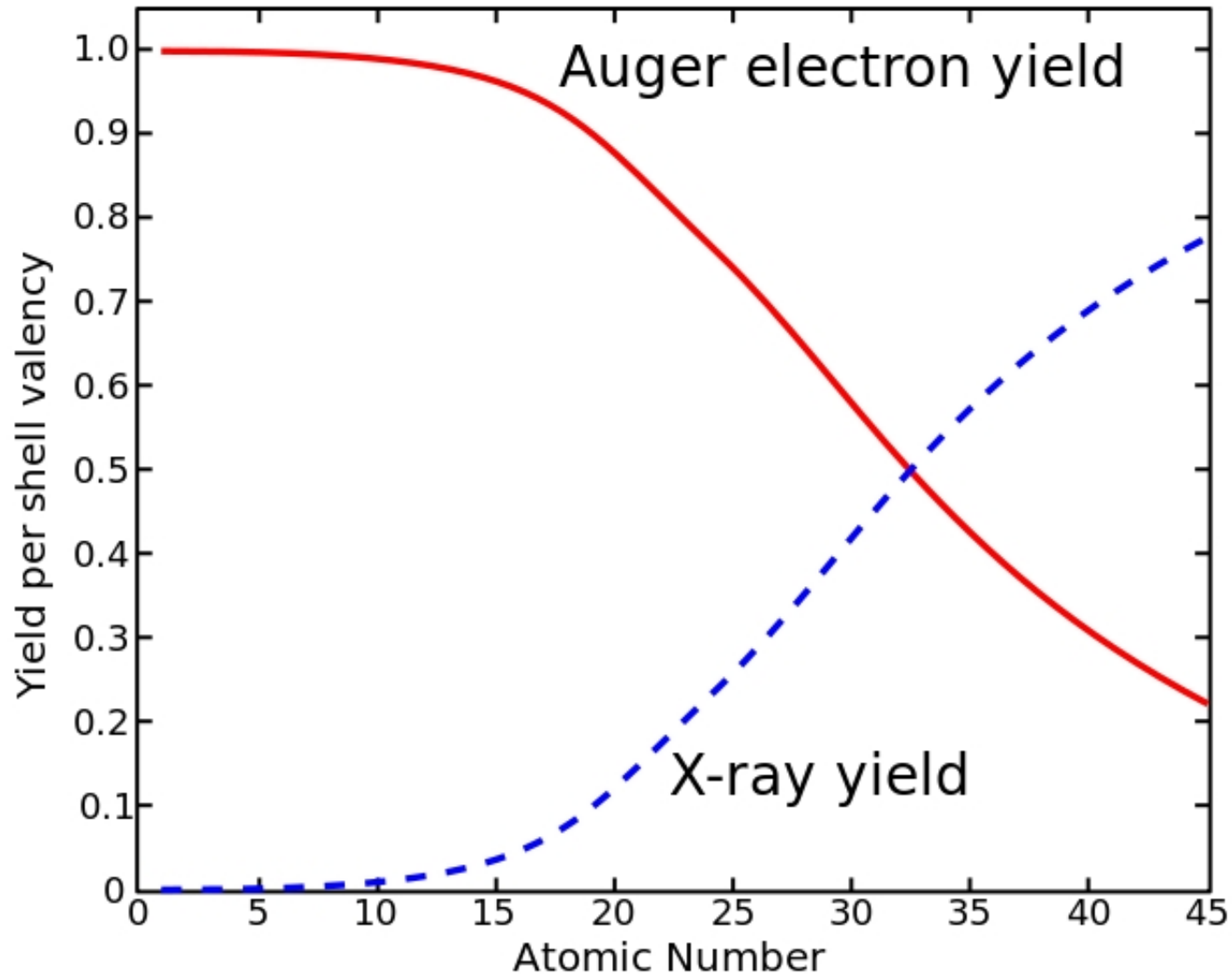
- The importance of the **inelastic scatter** to the practical analysis is connected to the appearance of scattered photons of lower energy than the incident photon beam (well separated for e.g. ^{241}Am when the ~60 keV elastic and ~49 keV inelastic scattering are clearly visible, see Fig.17)

The Escape Peak – another effect related to the detector efficiency

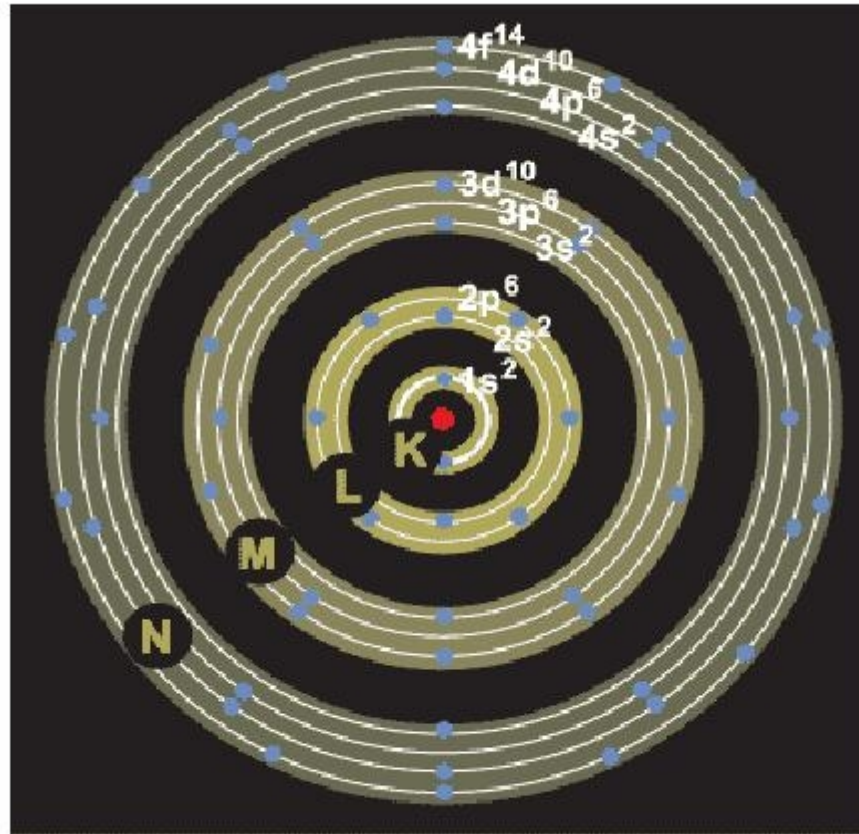
- The energy transfer in the detector is in part through photo-electric absorption, creating Si or Ge x-rays which are being reabsorbed.
- Si escape energy is E-1.74 keV
- Ge more serious problem E-9.97($K_{\alpha 1}$)



Schematic representation of the Auger Effect KLL



Fluorescence and Auger electron yields as a function of atomic number for K shell vacancies. Auger transitions (red curve) are more probable for lighter elements, while X-ray yield (dotted blue curve) becomes dominant at higher atomic numbers. Similar plots can be obtained for L and M shell transitions.



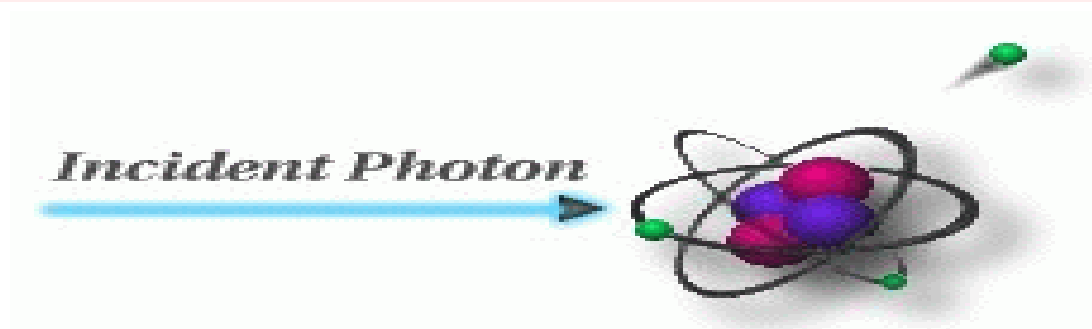
$$2(2l+1)$$

- 2 for s; $l=0$
- 6 for p; $l=1$
- 10 for d; $l=2$
- 14 for f; $l=3$

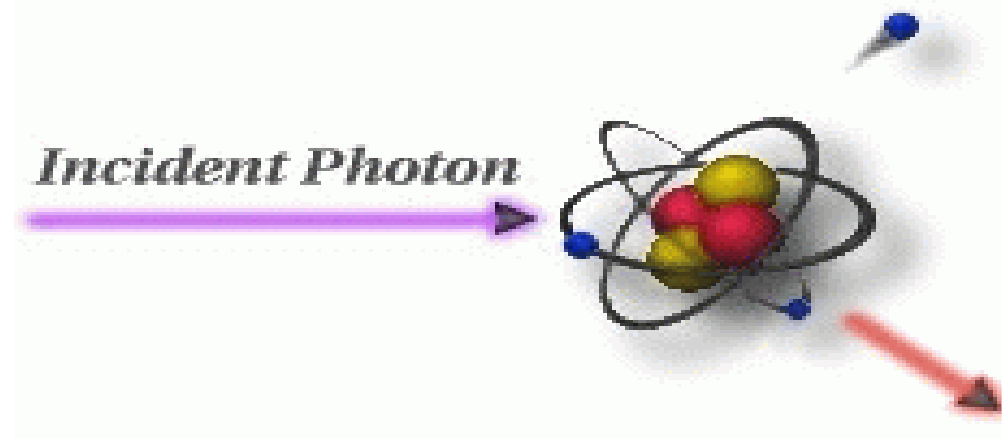


There are at least 4 modes through which electromagnetic radiations interact with atom: i.e., photoelectric absorption, Compton scattering (inelastic scattering), elastic scattering and pair production.

Photoelectric (PE) absorption of x-rays occurs when the x-ray photon is absorbed, resulting in the ejection of electrons from the outer shell of the atom, and hence the ionization of the atom. Subsequently, the ionized atom returns to the neutral state with the emission of an x-ray characteristic of the atom. This subsequent emission of lower energy photons is generally absorbed and does not contribute to (or hinder) the image making process. Photoelectron absorption is the dominant process for x-ray absorption up to energies of about 500 KeV. Photoelectron absorption is also dominant for atoms of high atomic numbers.

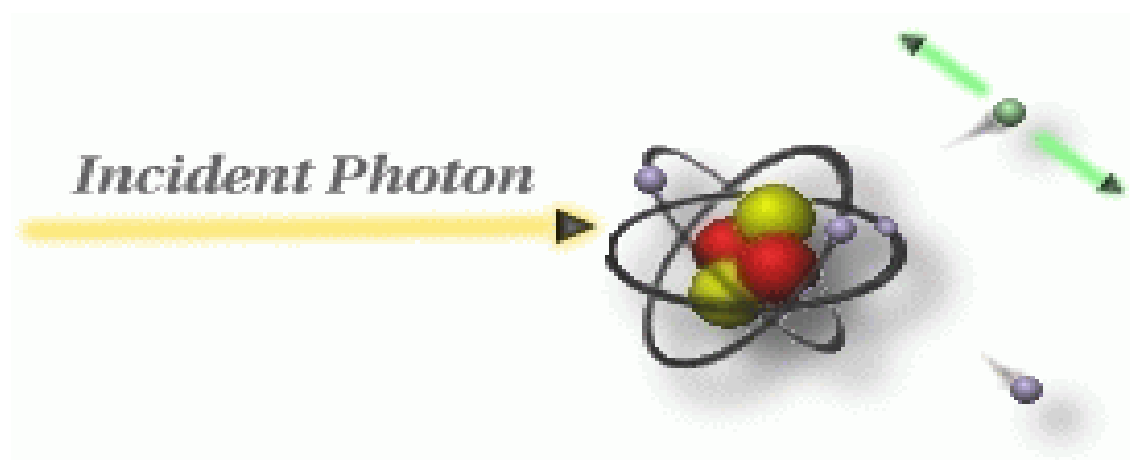


Compton scattering (C) occurs when the incident x-ray photon is deflected from its original path by an interaction with an electron. The electron gains energy and is ejected from its orbital position. The x-ray photon loses energy due to the interaction but continues to travel through the material along an altered path. Since the scattered x-ray photon has less energy, it, therefore, has a longer wavelength than the incident photon. The event is also known as incoherent scattering because the photon energy change resulting from an interaction is not always orderly and consistent. The energy shift depends on the angle of scattering and not on the nature of the scattering medium.



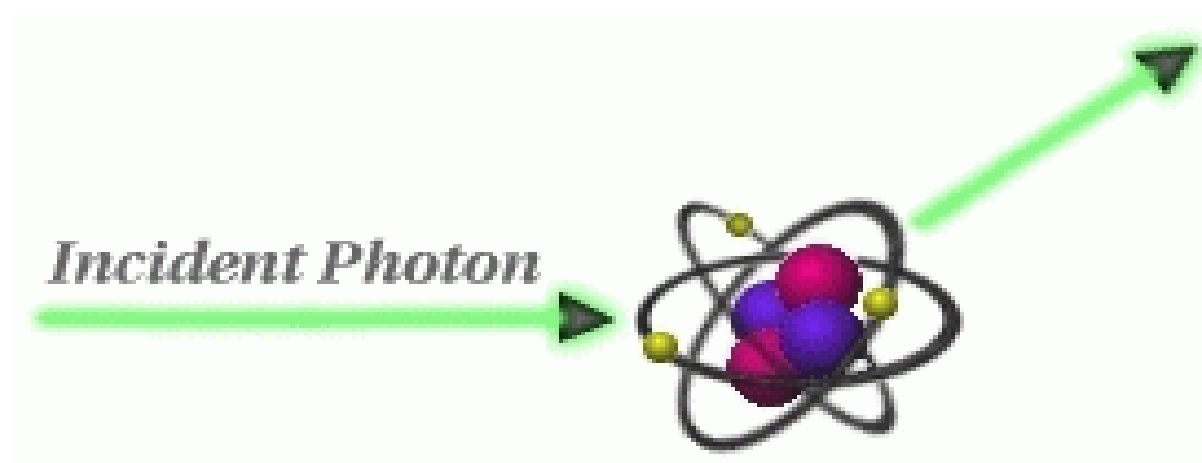
Compton-Incoherent or Inelastic $\sigma \sim Z$ (compton escape continuum -detector effect)

Pair production (PP) can occur when the x-ray photon energy is greater than 1.02 MeV, but really only becomes significant at energies around 10 MeV. Pair production occurs when an electron and positron are created with the annihilation of the x-ray photon. Positrons are very short lived and disappear (positron annihilation) with the formation of two photons of 0.51 MeV energy. Pair production is of particular importance when high-energy photons pass through materials of a high atomic number.

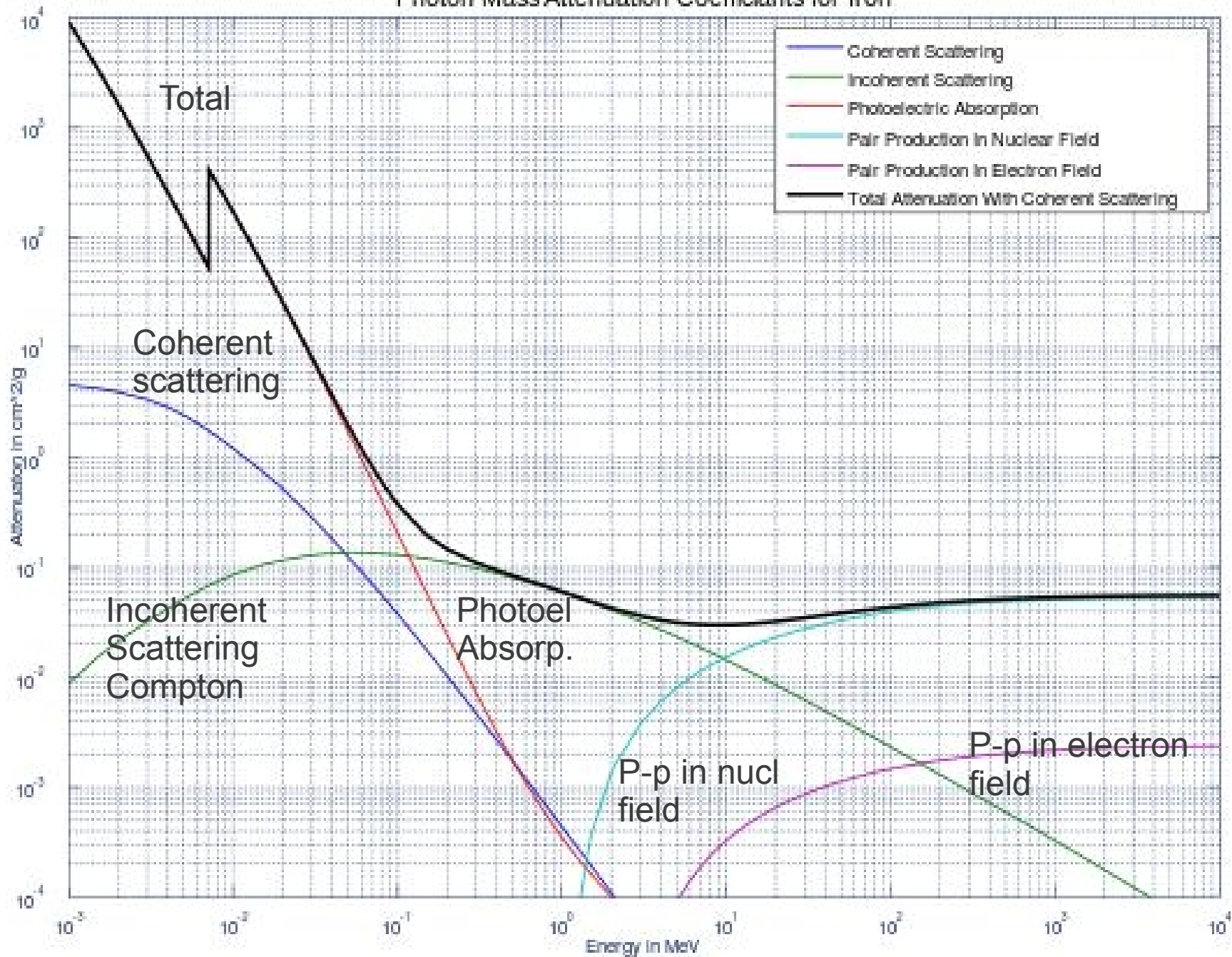


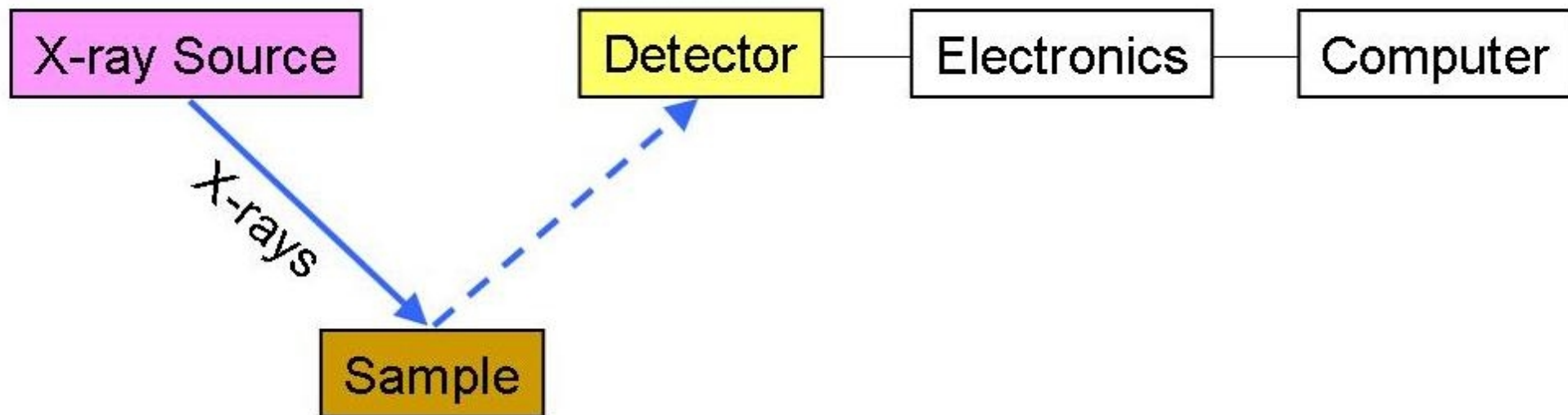
Elastic scattering of γ -rays and X-rays by atoms

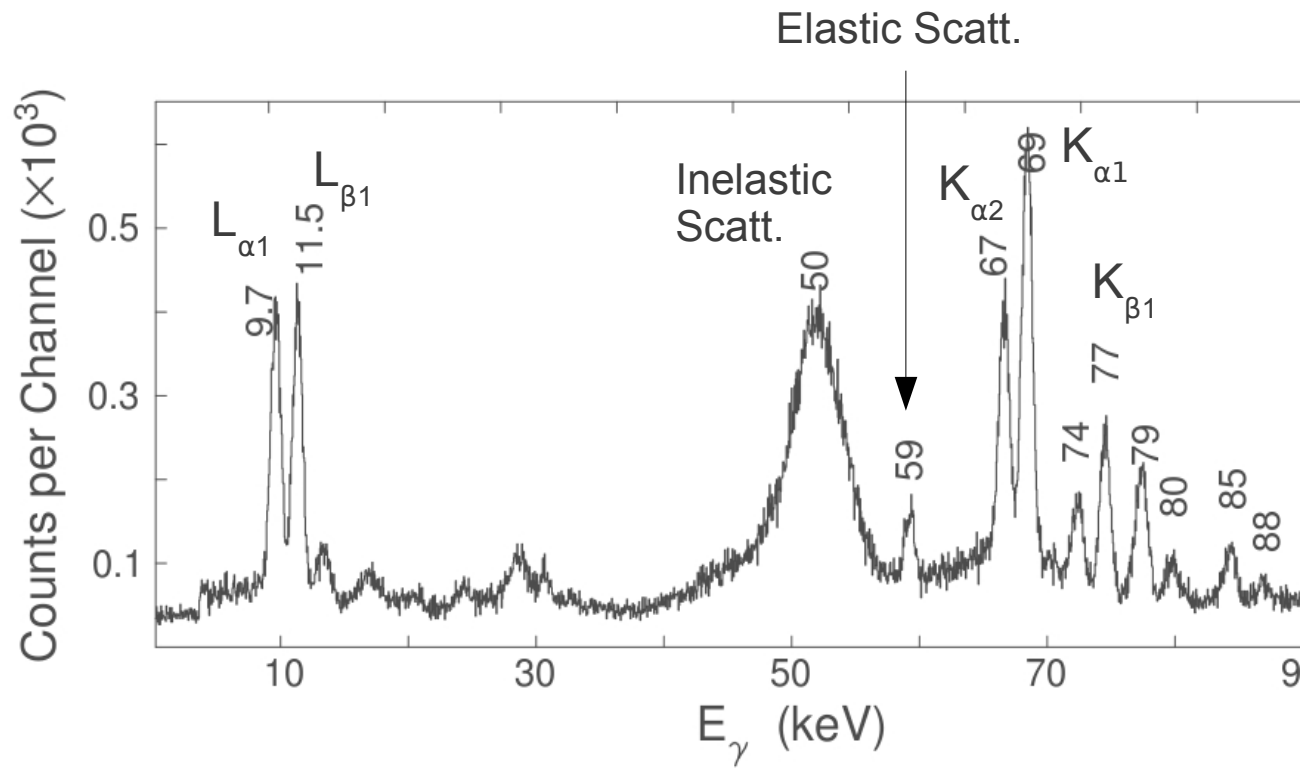
Thomson scattering (R), also known as Rayleigh, **coherent**, or classical scattering, occurs when the x-ray photon interacts with the whole atom so that the photon is scattered with no change in internal energy to the scattering atom, nor to the x-ray photon. Thomson scattering is never more than a minor contributor to the absorption coefficient. The scattering occurs without the loss of energy. Scattering is mainly in the forward direction.



Photon Mass Attenuation Coefficients for Iron







Typical scattering spectrum of 59.9 keV photons by the Au target through $\sim 130^\circ$. Most intensive K and L X-ray energies of Au element are marked.

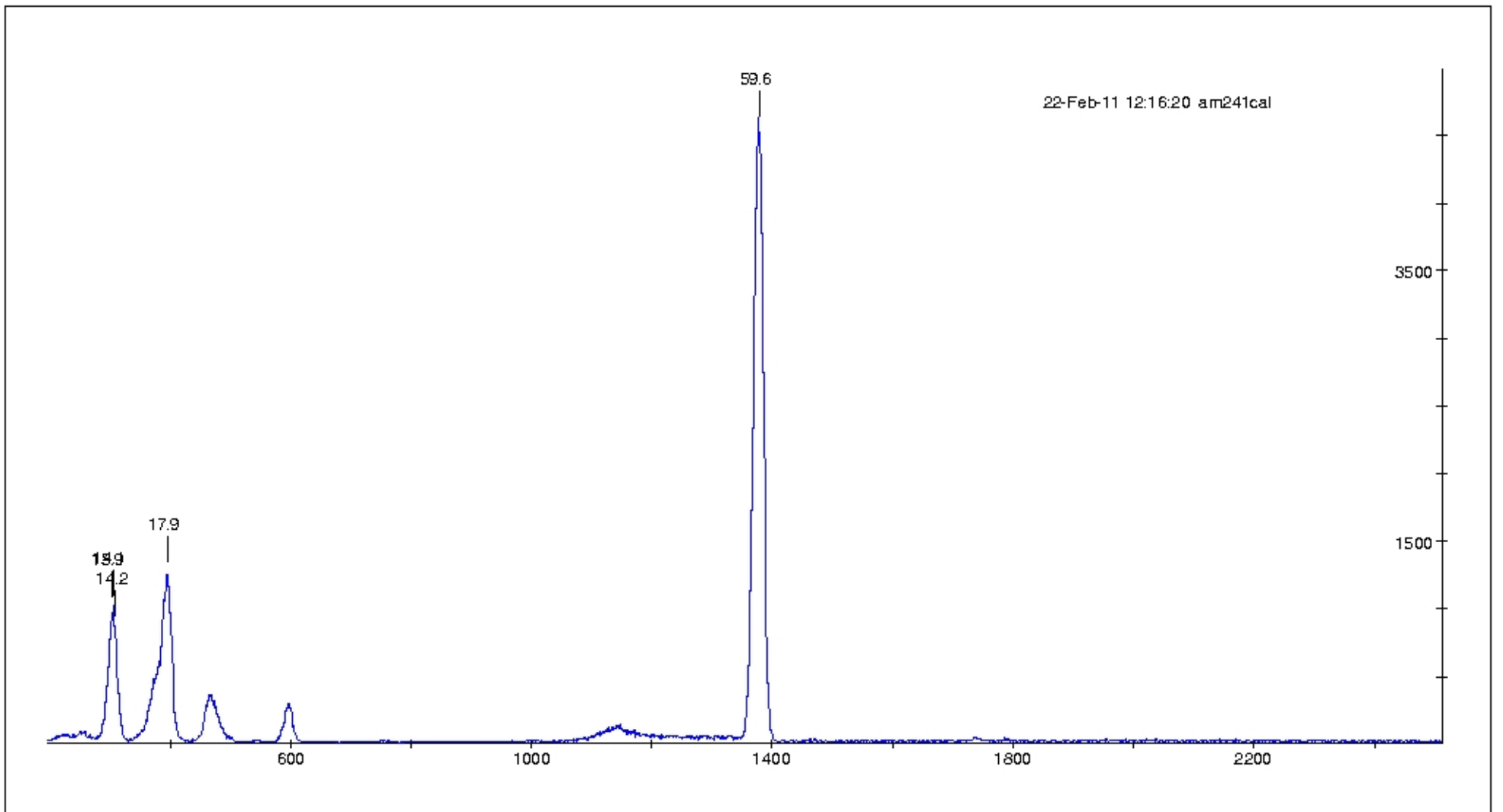


Table 7a. X-ray Energies and Intensities (per 100 K-Shell Vacancies)

	5 B	6 C	7 N	8 O	9 F	10 Ne	11 Na	12 Mg	13 Al	14 Si	15 P	16 S	17 Cl	18 Ar	19 K
$K_{\alpha 1}$	0.183 0.115	0.277 0.198	0.392 0.3514	0.525 0.5522	0.677 0.94	0.849 1.2012	1.041 1.5316	1.254 2.02	1.487 2.63	1.740 3.33	2.010 4.14	2.308 5.05	2.622 6.15	2.957 7.37	3.314 8.59
$K_{\alpha 2}$	0.183 0.05628	0.277 0.094	0.392 0.177	0.525 0.2811	0.677 0.4317	0.848 0.808	1.041 0.778	1.254 1.0010	1.486 1.2913	1.739 1.6417	2.009 2.0421	2.307 2.4925	2.621 3.03	2.955 3.64	3.311 4.34
$K_{\beta 1}$									1.554 0.015516	1.836 0.0565	2.136 0.12212	2.464 0.22923	2.816 0.384	3.190 0.585	3.590 0.798
$K_{\beta 3}$									1.554 0.00798	1.836 0.0283	2.136 0.0625	2.464 0.11612	2.816 0.19220	3.190 0.303	3.590 0.404
$L_{\beta 1}$														0.251 0.0113	0.296 0.0134
$L_{\beta 3}$														0.310 0.003813	0.359 0.005017
$L_{\beta 4}$														0.310 0.00249	0.359 0.00105
	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se
$K_{\alpha 1}$	3.692 9.84	4.091 11.35	4.511 12.85	4.952 14.57	5.415 16.47	5.899 18.38	6.404 20.29	6.930 22.110	7.478 24.011	8.048 26.012	8.639 28.010	9.252 29.811	9.886 31.311	10.544 32.712	11.222 34.112
$K_{\alpha 2}$	3.688 4.9322	4.086 5.6825	4.505 6.43	4.945 7.33	5.405 8.34	5.888 9.34	6.391 10.25	6.915 11.25	7.461 12.25	8.028 13.35	8.616 14.35	9.225 15.25	9.855 16.15	10.508 16.85	11.182 17.65
$K_{\beta 1}$	4.013 1.025	4.461 1.225	4.932 1.425	5.427 1.647	5.947 1.848	6.490 2.1410	7.058 2.4011	7.649 2.6512	8.265 2.8813	8.905 3.1014	9.572 3.3912	10.264 3.7013	10.982 3.9814	11.726 4.2515	12.496 4.5416
$K_{\beta 2}$												10.366 0.031411	11.101 0.0974	11.864 0.1947	12.652 0.3233
$K_{\beta 3}$	4.013 0.51928	4.461 0.623	4.932 0.723	5.427 0.844	5.947 0.944	6.490 1.095	7.058 1.235	7.649 1.365	8.265 1.487	8.905 1.597	9.572 1.745	10.260 1.907	10.975 2.057	11.720 2.198	12.490 2.348
$K_{\beta 5}$							7.108 0.001277	7.706 0.0018811	8.329 0.0026415	8.977 0.0036521	9.651 0.0050425	10.350 0.00633	11.074 0.00784	11.826 0.00955	12.601 0.01165
$L_{\alpha 1}$		0.396 0.0287	0.452 0.06316	0.511 0.123	0.572 0.195	0.637 0.267	0.704 0.338	0.776 0.4110	0.851 0.5013	0.929 0.6015	1.012 0.6513	1.098 0.7014	1.188 0.8116	1.282 0.8717	1.379 0.9820
$L_{\alpha 2}$		0.396 0.00287	0.452 0.007018	0.511 0.0133	0.572 0.0215	0.637 0.0297	0.704 0.0379	0.776 0.04511	0.851 0.05614	0.929 0.06617	1.012 0.07215	1.098 0.07716	1.188 0.09018	1.282 0.09619	1.379 0.10822
$L_{\beta 1}$	0.350 0.0164	0.400 0.0205	0.458 0.05012	0.518 0.09624	0.581 0.154	0.648 0.205	0.717 0.255	0.791 0.318	0.868 0.349	0.949 0.3910	1.035 0.4211	1.125 0.4612	1.219 0.4912	1.317 0.5213	1.420 0.5815
$L_{\beta 3}$	0.412 0.006219	0.468 0.007523	0.529 0.0093	0.590 0.0103	0.652 0.0124	0.720 0.0144	0.792 0.0165	0.866 0.0185	0.940 0.0205	1.022 0.0216	1.107 0.0237	1.195 0.0247	1.294 0.0257	1.386 0.0278	1.492 0.0299
$L_{\beta 4}$	0.412 0.003912	0.468 0.004815	0.529 0.005617	0.590 0.006720	0.652 0.007924	0.720 0.0093	0.792 0.0103	0.866 0.0124	0.940 0.0134	1.022 0.0144	1.107 0.0155	1.191 0.0165	1.286 0.0165	1.380 0.0185	1.486 0.0195
$L_{\beta 6}$		0.402 0.00174	0.456 0.00185	0.513 0.00225		0.640 0.00235	0.708 0.00225	0.779 0.00225	0.855 0.00225	0.940 0.00225	1.022 0.00214	1.107 0.00275	1.191 0.00337	1.286 0.00388	1.380 0.00459
$L_{\gamma 3}$												1.297 0.00124	1.412 0.004213	1.524 0.004715	1.648 0.005116
L_{η}		0.353 0.0205	0.401 0.0225	0.454 0.0267	0.510 0.0255	0.568 0.0267	0.628 0.0287	0.693 0.0287	0.760 0.0267	0.831 0.0287	0.907 0.0297	0.984 0.0308	1.068 0.0318	1.155 0.0318	1.245 0.0349
L_{ζ}		0.348 0.0267	0.395 0.0298	0.446 0.0349	0.500 0.0339	0.556 0.03810	0.615 0.04011	0.678 0.04311	0.743 0.04512	0.811 0.04813	0.884 0.04710	0.957 0.04810	1.037 0.05211	1.120 0.05311	1.204 0.05612
	35 Br	36 Kr	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In
$K_{\alpha 1}$	11.924 35.613	12.651 36.813	13.395 38.014	14.165 39.114	14.958 40.114	15.775 41.012	16.615 41.812	17.479 42.612	18.367 43.312	19.279 44.012	20.216 44.613	21.177 45.113	22.163 45.613	23.174 46.113	24.210 45.313
$K_{\alpha 2}$	11.878 18.47	12.598 19.07	13.336 19.77	14.098 20.37	14.883 20.98	15.691 21.45	16.521 21.95	17.374 22.45	18.251 22.85	19.150 23.27	20.074 23.57	21.020 23.97	22.000 24.27	22.984 24.57	24.002 24.57
$K_{\alpha 3}$													21.708 0.001004	22.693 0.001154	23.702 0.001355
$K_{\beta 1}$	13.292 4.8417	14.111 5.1219	14.961 5.3919	15.836 5.6320	16.738 5.8921	17.667 6.1517	18.623 6.3518	19.607 6.6119	20.619 6.8019	21.657 6.9920	22.724 7.1820	23.819 7.3521	24.943 7.5221	26.095 7.6922	27.276 7.8522
$K_{\beta 2}$	13.469 0.48419	14.311 0.67624	15.185 0.853	16.085 1.004	17.013 1.134	17.969 1.254	18.952 1.334	19.965 1.454	21.005 1.544	22.074 1.645	23.172 1.725	24.299 1.795	25.455 1.885	26.644 1.985	27.863 2.095
$K_{\beta 3}$	13.284 2.509	14.104 2.6410	14.952 2.7810	15.825 2.9110	16.726 3.0411	17.653 3.179	18.607 3.289	19.590 3.4110	20.599 3.5110	21.634 3.6110	22.699 3.7110	23.791 3.8111	24.912 3.9011	26.060 3.9911	27.238 4.0712
$K_{\beta 4}$							18.982 0.00105	19.998 0.00157	21.042 0.002311	22.115 0.003216	23.217 0.004321	24.349 0.0063	25.511 0.0073	26.702 0.0084	27.924 0.0105
$K_{\beta 5}$	13.404 0.01397	14.231 0.01628	15.089 0.01869	15.971 0.021511	16.880 0.024412	17.816 0.027512	18.780 0.030514	19.771 0.034115	20.789 0.037717	21.836 0.041819	22.911 0.044620	24.013 0.049622	25.144 0.054725	26.304 0.0603	27.493 0.0653
$KO_{2,3}$															27.939 0.017018
$L_{\alpha 1}$	1.481 1.0922	1.581 1.2024	1.694 1.33	1.806 1.43	1.923 1.53	2.042 1.6625	2.166 1.83	2.293 1.93	2.424 2.03	2.558 2.13	2.697 2.33	2.839 2.44	2.984 2.54	3.134 2.64	3.287 2.84